

DARK-MATTER

ANTI-MATTER

AMS-02

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Trento University & TIFPA



Big Questions

What the universe is made of ?

How the Universe change with time ?

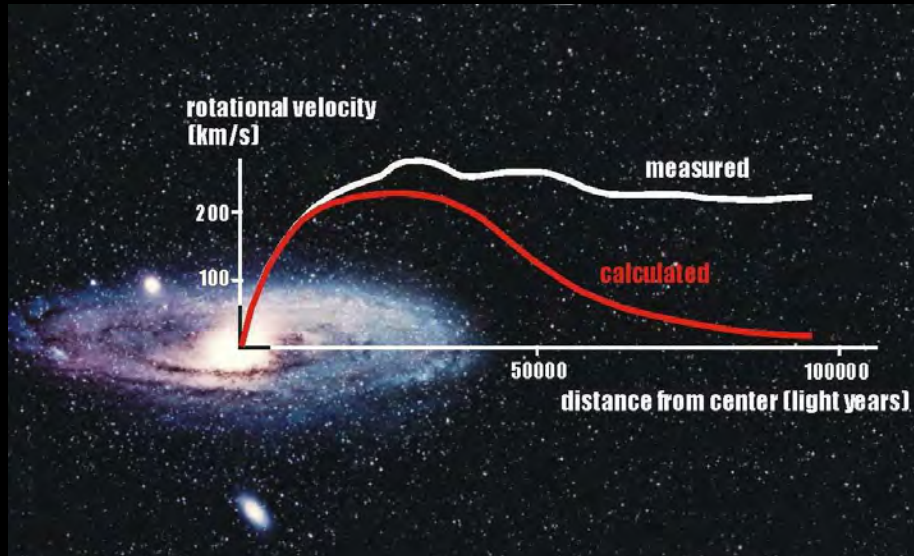
Why matter is organized in stars, galaxies, ...?

How fundamental particle physics is related to the large scale?

Dark Matter history

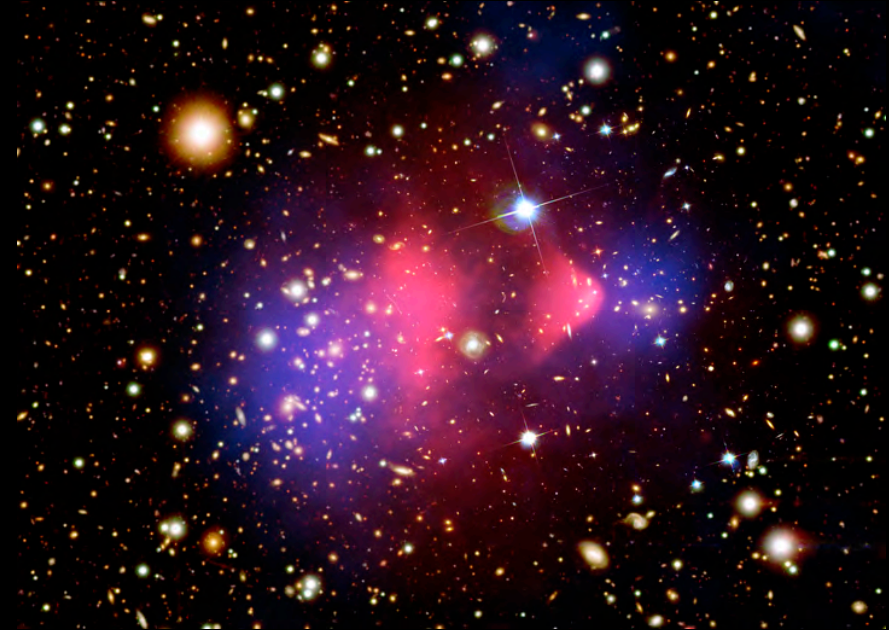
**We know about the Dark Matter existence
from his gravitational effects**

Galaxy rotation curves

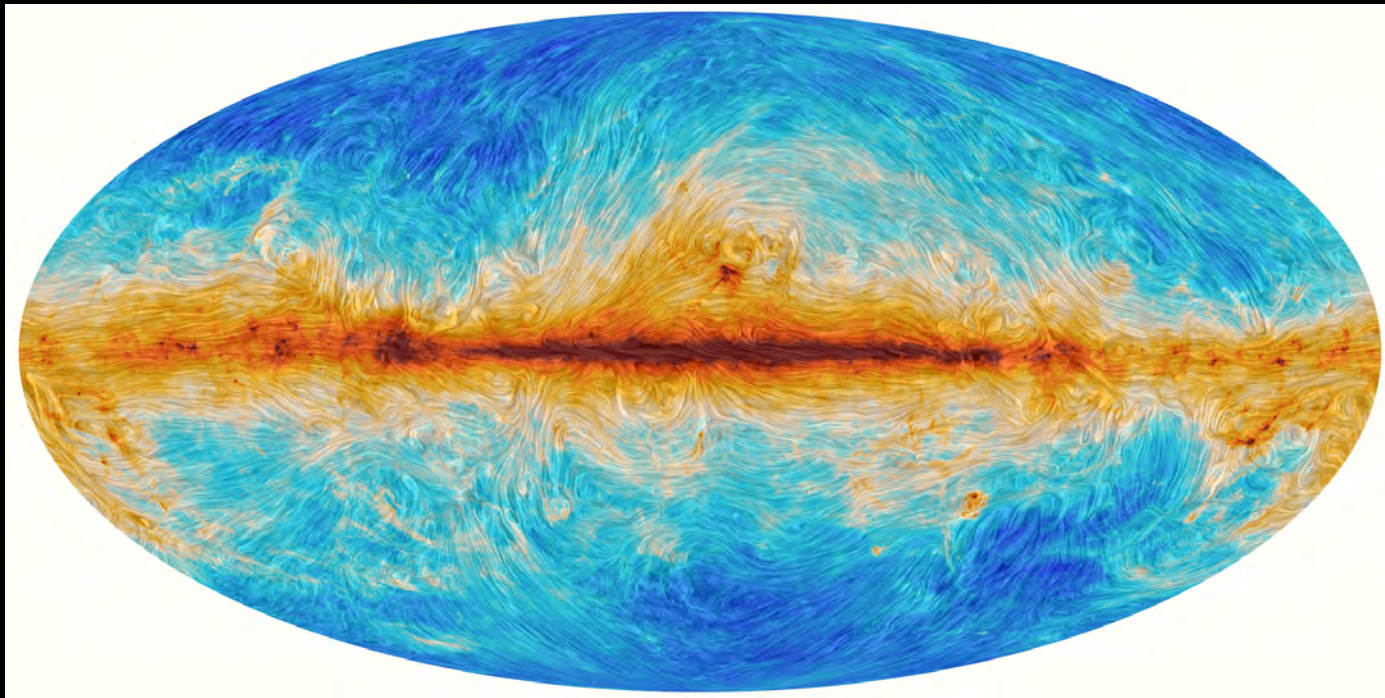


Relative motion of Galaxies

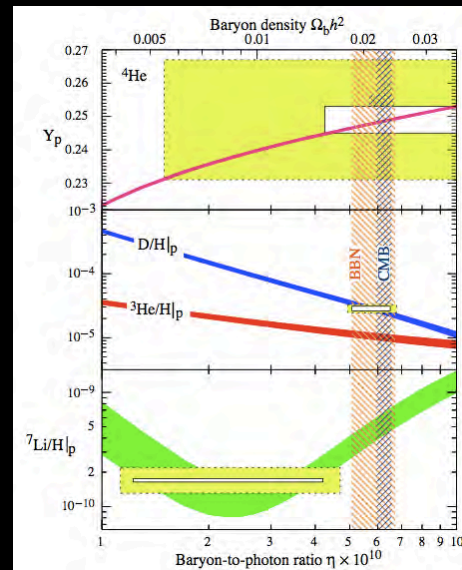
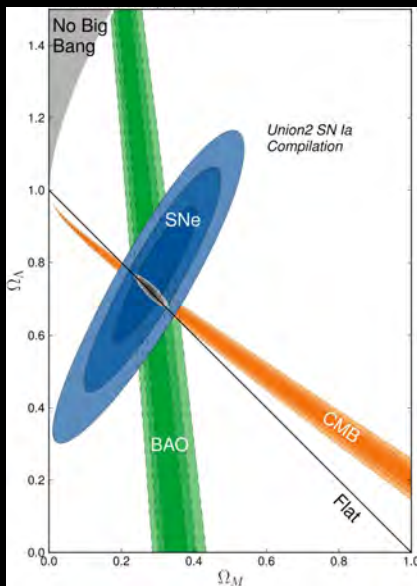
The Bullet cluster



The PLANK CMB sky

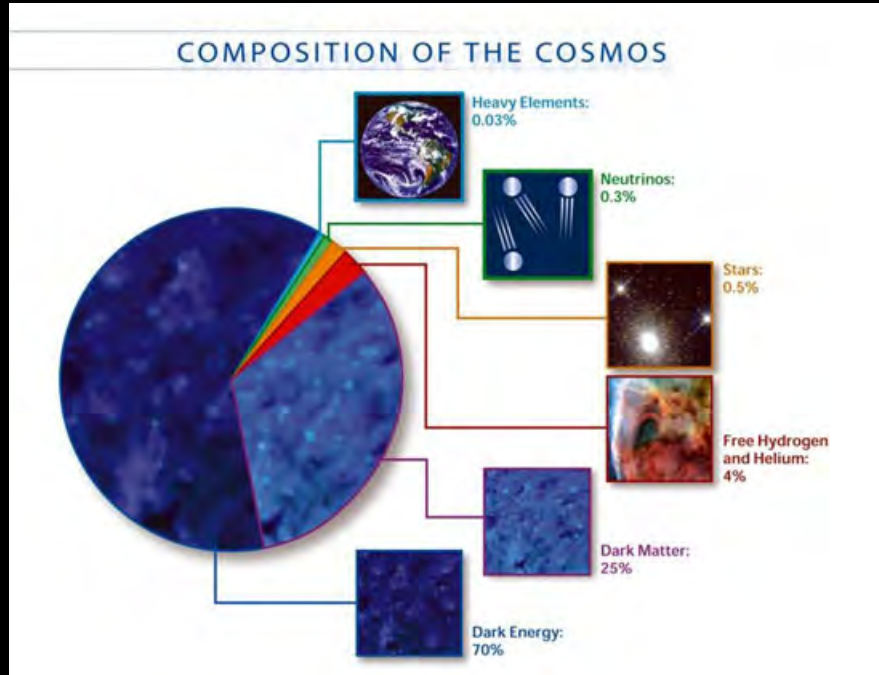


The Cosmic Equation



$$\Omega_T = \Omega_\Lambda + \Omega_{DM} + \Omega_{LM}$$

$$\Omega_{NB} + \Omega_B$$



What Dark Energy is ?

What Dark Matter is made of ?

Is Anti-Matter really gone ?

the Cosmic Rays

A flux of charged particle from the galaxy that constantly irradiate the Earth

Their abundance versus the energy follows a power law
law : $F = E^{-k}$

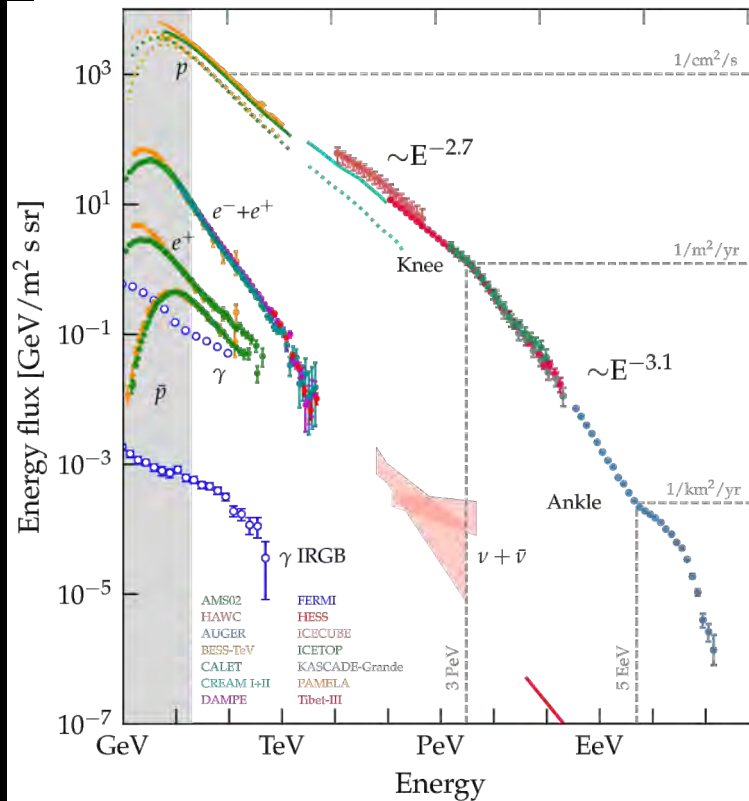
$$\log F = -k \log E$$

The energy spectrum cover 12 orders of magnitude for an intensity that spans over ~ 30 orders of magnitude

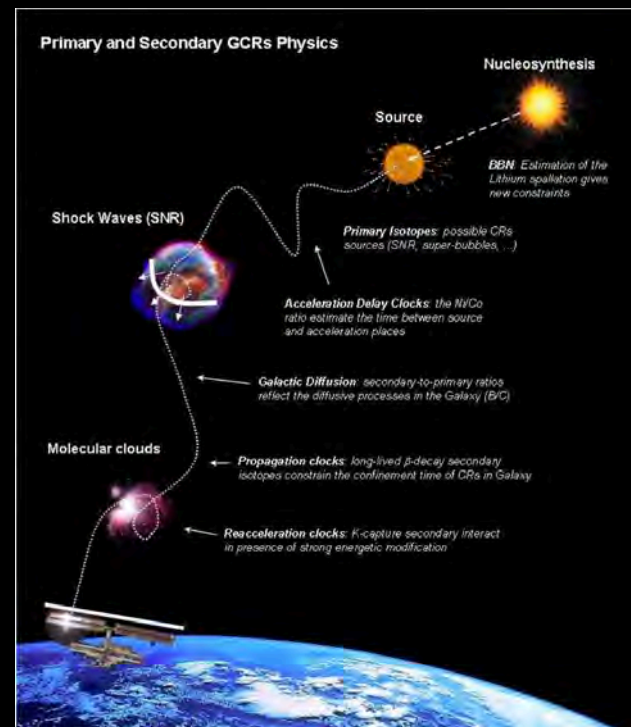
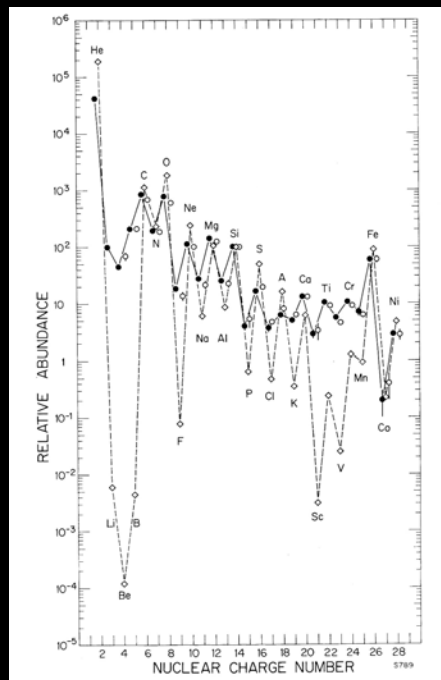
Composition:

- 89% protons
- 9% He nuclei
- $\sim 1\%$ electrons
- $\sim 1\%$ $Z > 2$ Nuclei
- $< 0.1\%$ positrons & anti-protons

C. Evoli, <http://doi.org/10.5281/zenodo.2360277> (2018)



CR carry information about our galaxy



AMS-02: Alpha Magnetic Spectrometer

Launch 16/5/2011 (Endavour)

Construction 1999-2010

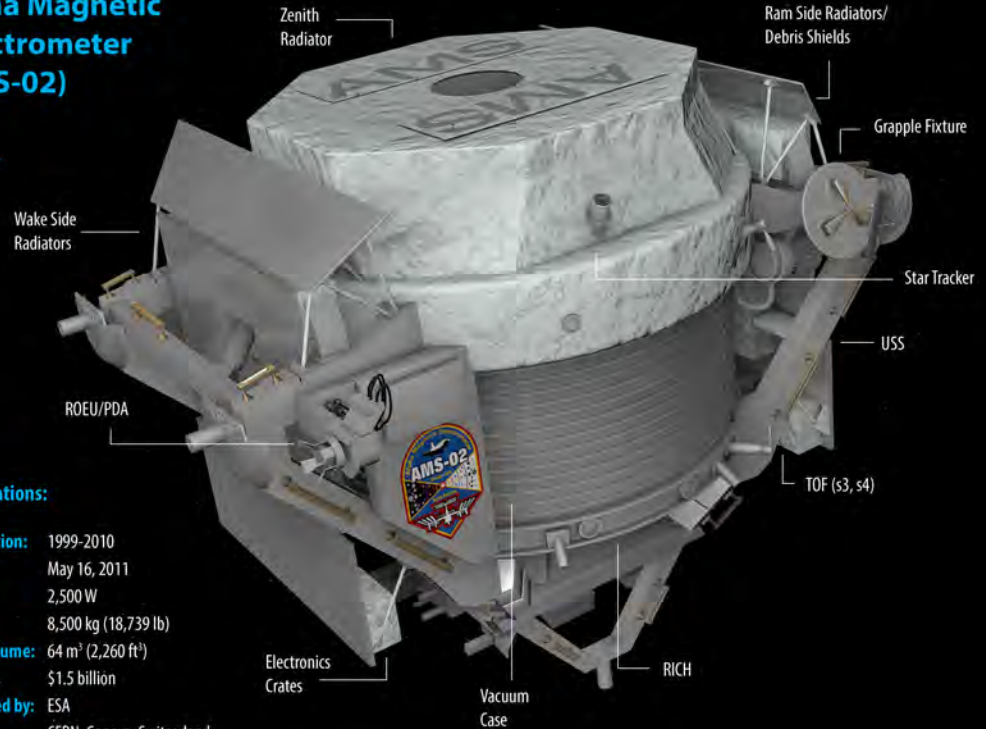
Dimensions $3 \times 4 \times 5 \text{ m}^3$

Weight 8.5 t

Power 2500 W

Alpha Magnetic Spectrometer (AMS-02)

Port view



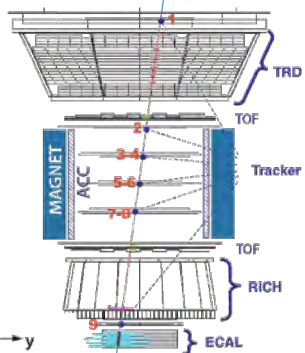
Specifications:

Construction: 1999-2010
Launch: May 16, 2011
Power: 2,500 W
Mass: 8,500 kg (18,739 lb)
Press. Volume: 64 m^3 (2,260 ft^3)
Cost: \$1.5 billion
Contracted by: ESA
Built by: CERN, Geneva, Switzerland



AMS-02 construction: EVA point of view

Tracker
planes



1



TRD

2



TOF

3-4



MAGNET

5-6

ACC



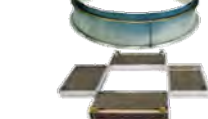
7-8



TOF



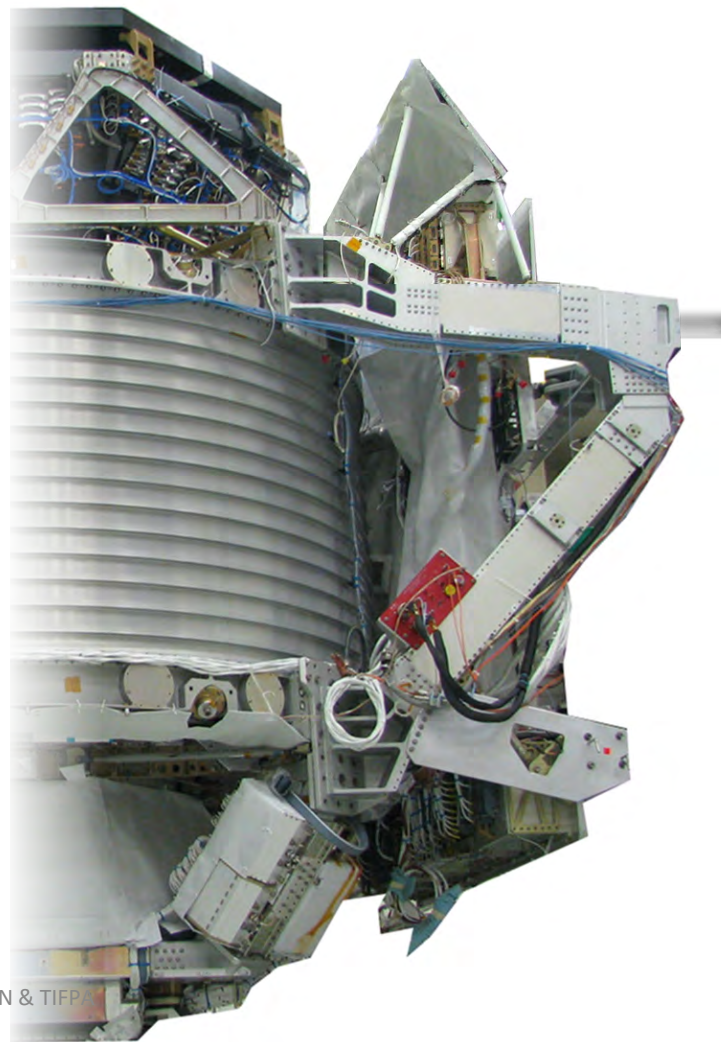
RICH



9

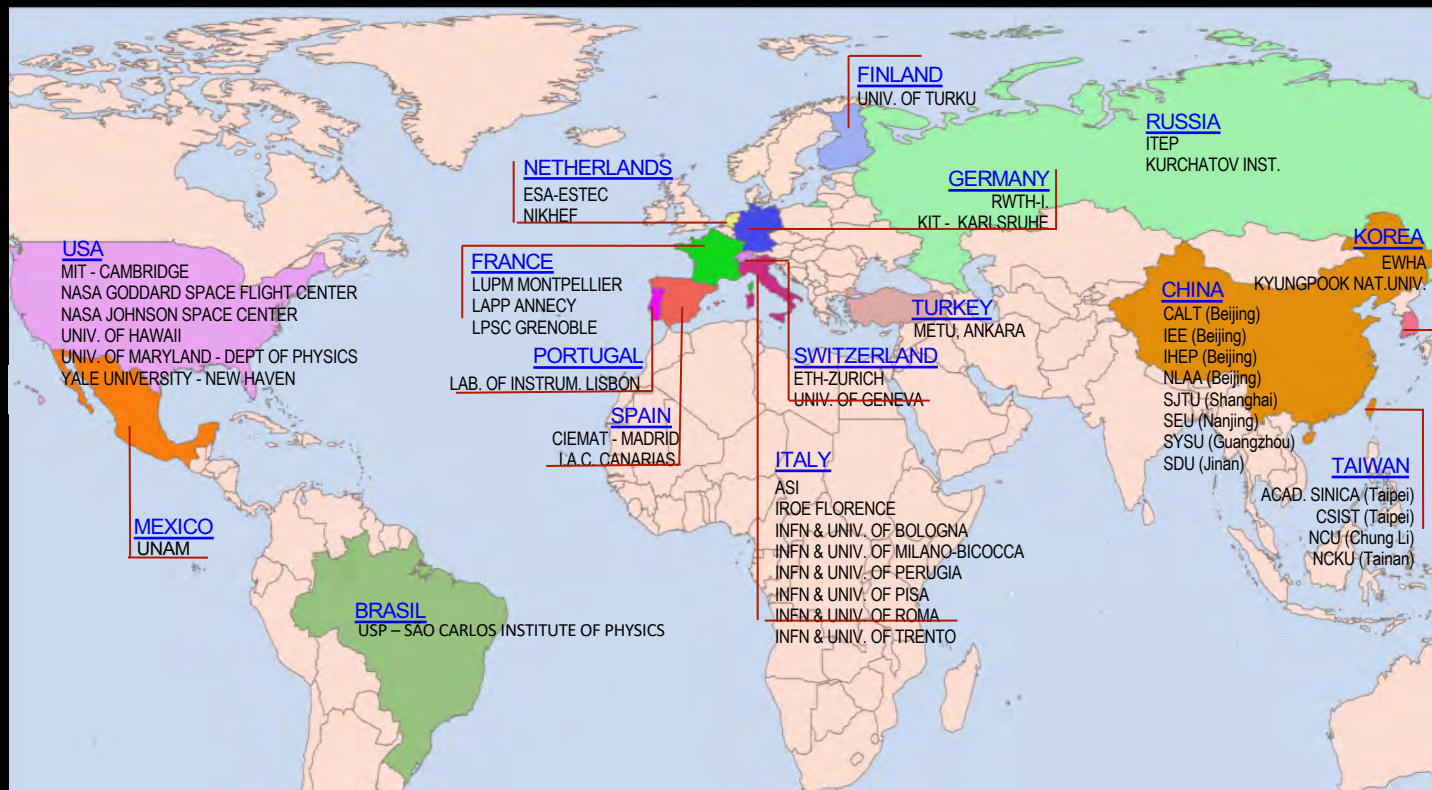


ECAL



AMS International Collaboration

16 Countries, 60 Institutes and 600 Physicists



AMS-02 is controlled 24/7 from CERN, Geneva, Switzerland



AMS at CERN

AMS-02 ITALIA



Università e INFN

Bologna

Milano "Bicocca"

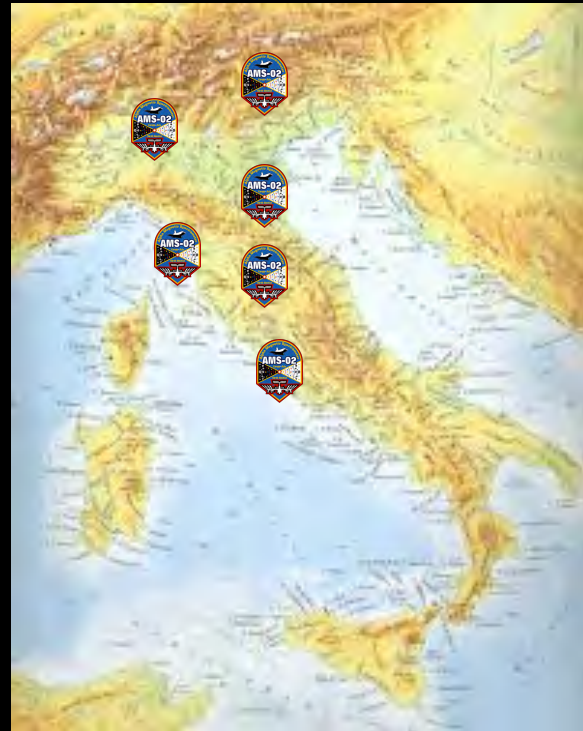
Perugia

Pisa

Roma "La Sapienza"

Roma "Tor Vergata"

Trento





AMS-02 in Trento

Francesco
Nozzoli



Paolo Zuccon



Roberto Battiston



William J. Burger



Francesco
Dimiccoli



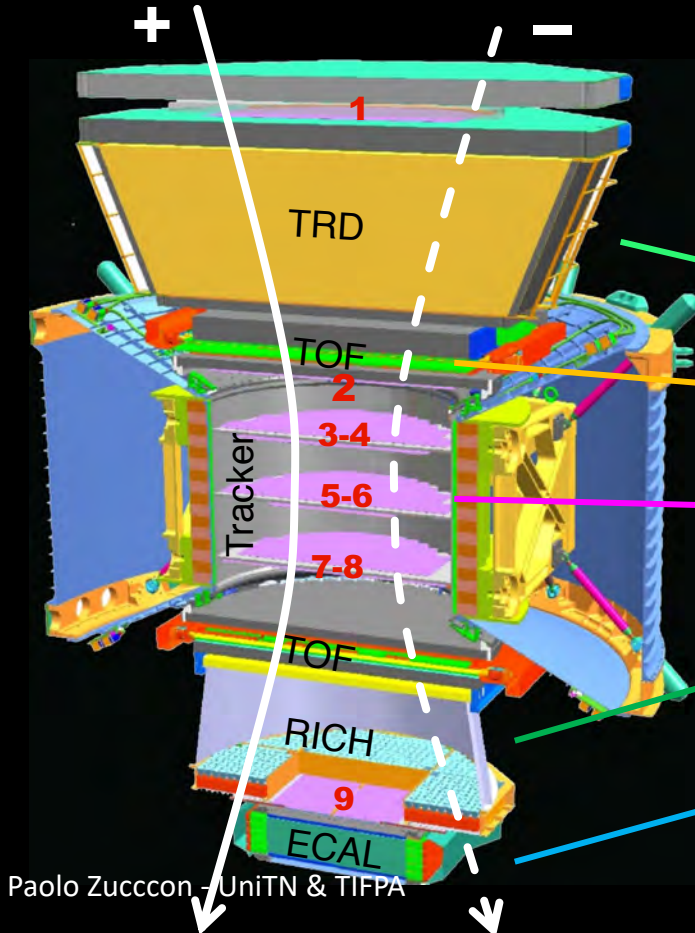
Cinzia Cernetti



Ignazio Lazizzera



AMS identifies particles, from the different signals, released in the 5 sub-detectors.



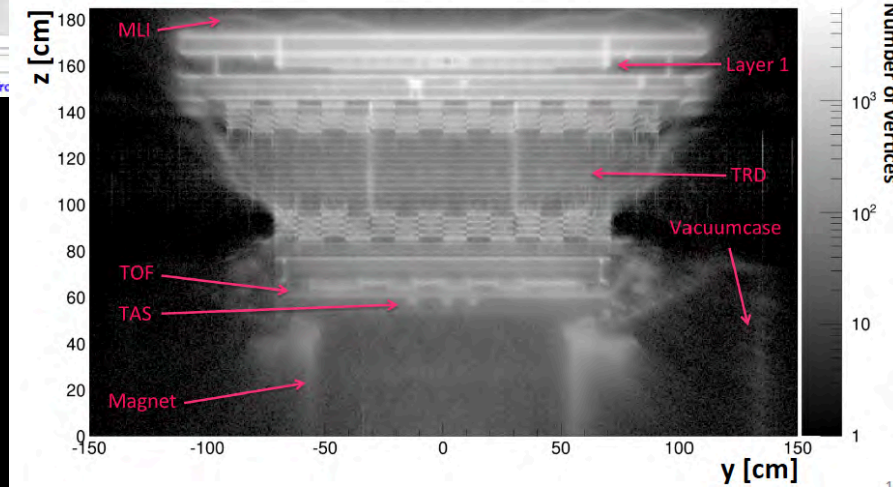
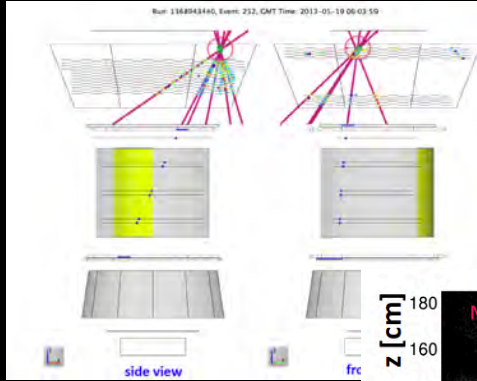
Matter

Anti-matter

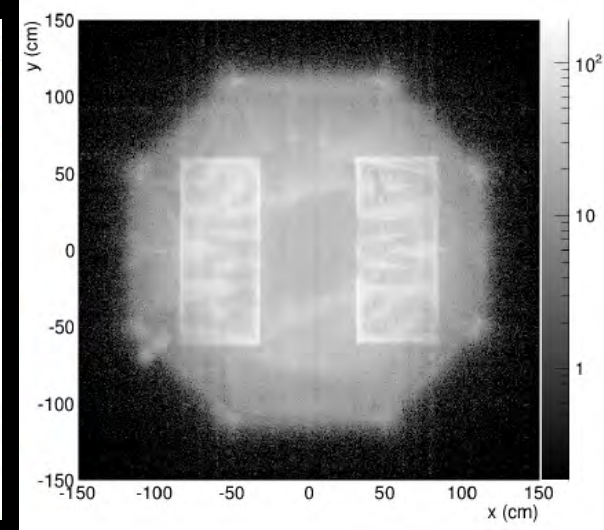
	e^-	P	Fe	e^+	$\bar{P}$	$\bar{He}$
TRD						
TOF						
Tracker + Magnet						
RICH						
ECAL						

AMS “tomography” using rare nuclear interaction events

16

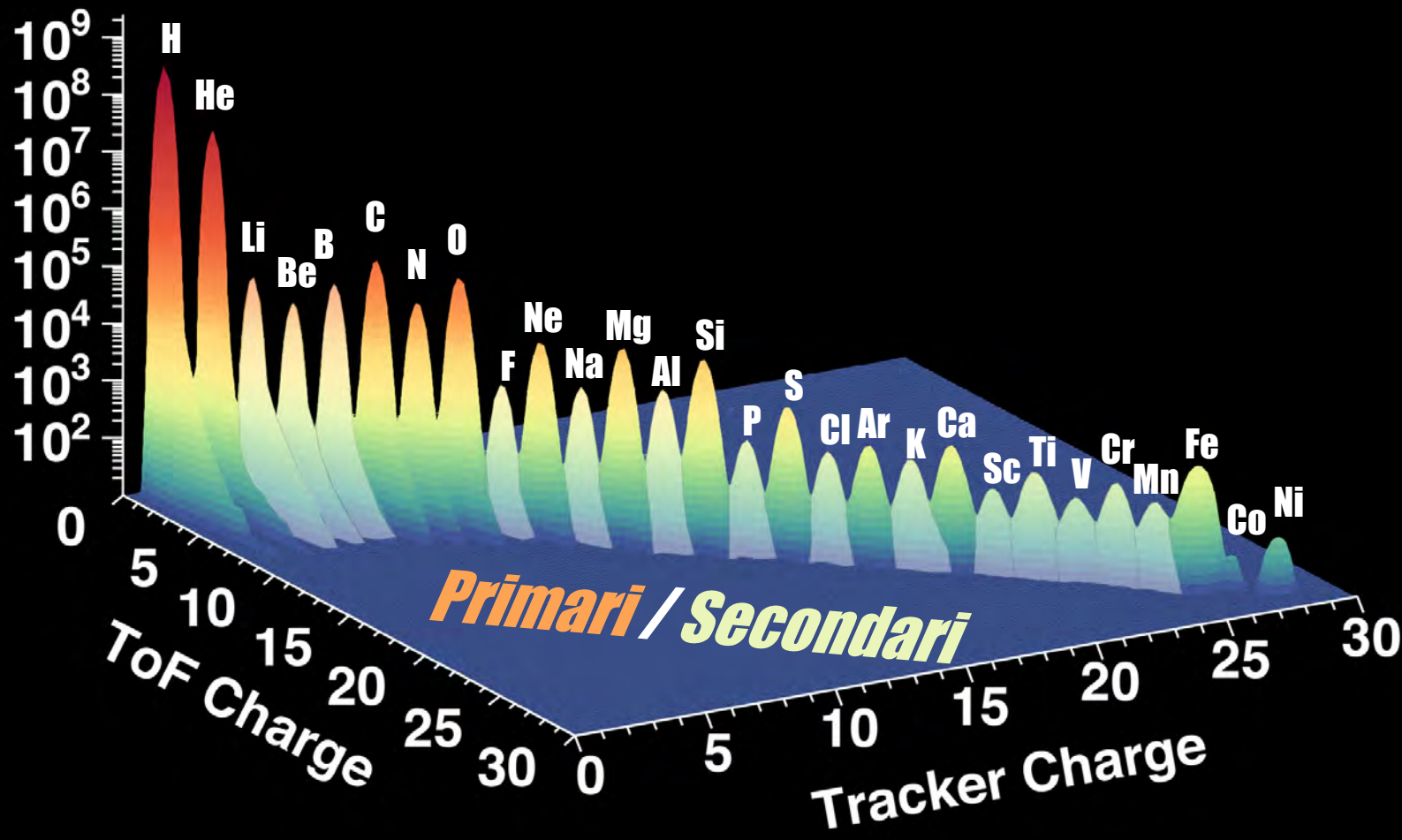


$Z=178.5$ cm



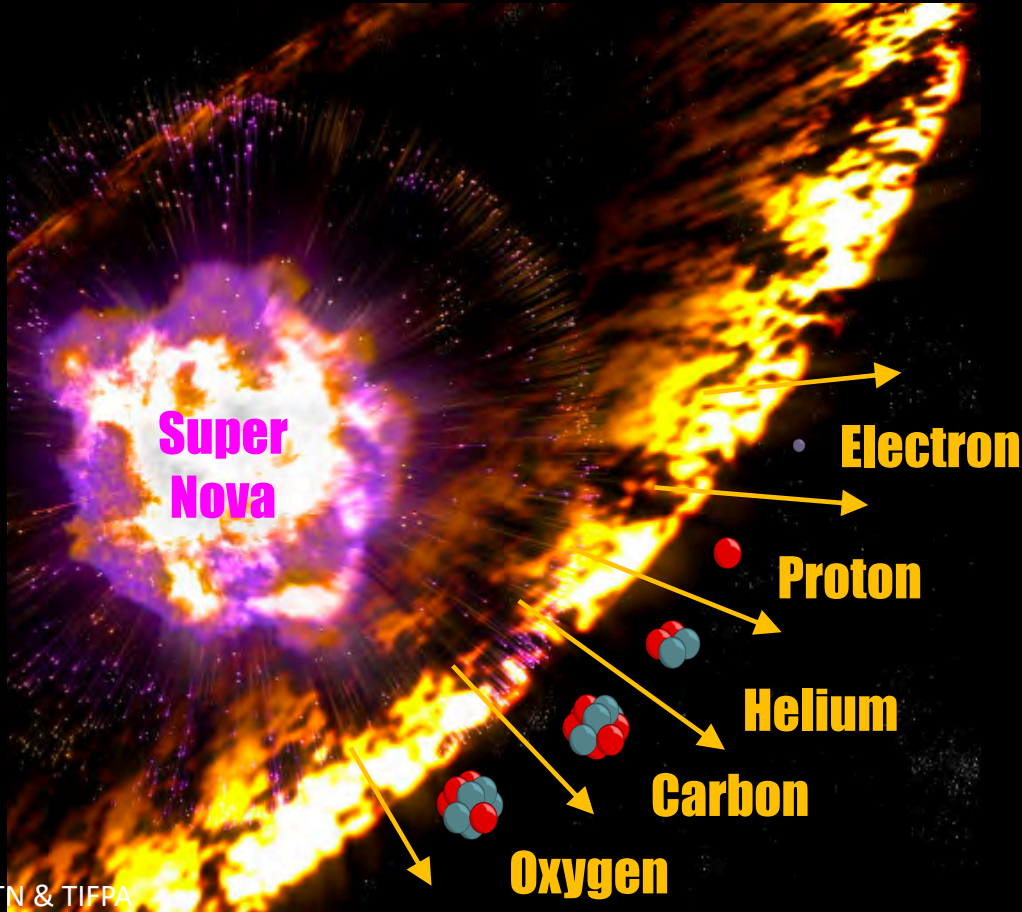
The gray scale is proportional the
the number of found vertices

Periodic table of elements with AMS-02

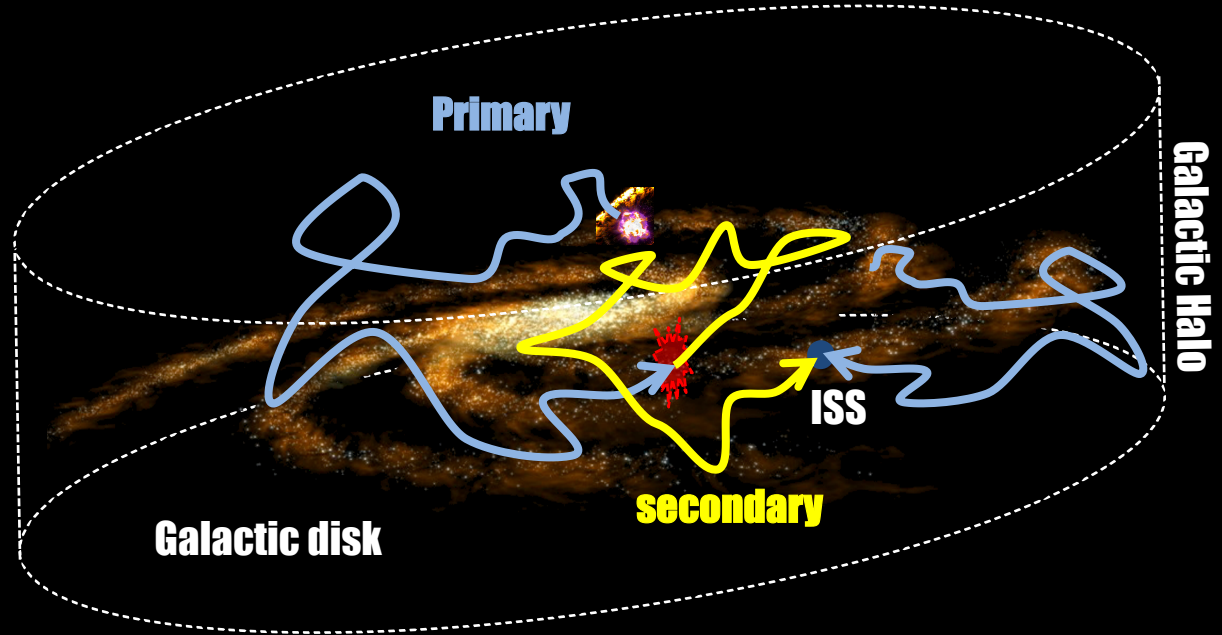


Primary Cosmic Rays

18



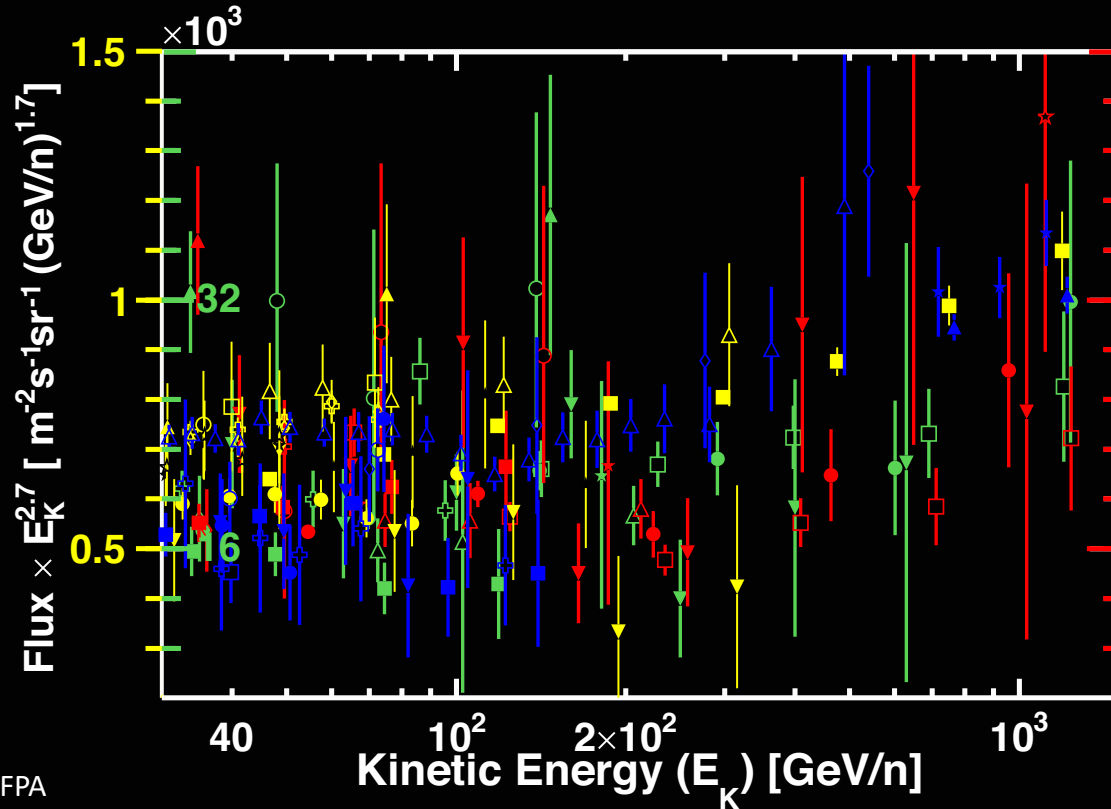
Secondary Cosmic Rays



**Primary cosmic rays are produced from astrophysics sources,
Secondary cosmic rays are produced in collisions with the inter stellar medium.**

Primaries before AMS-02

Helium
Carbon
Oxygen

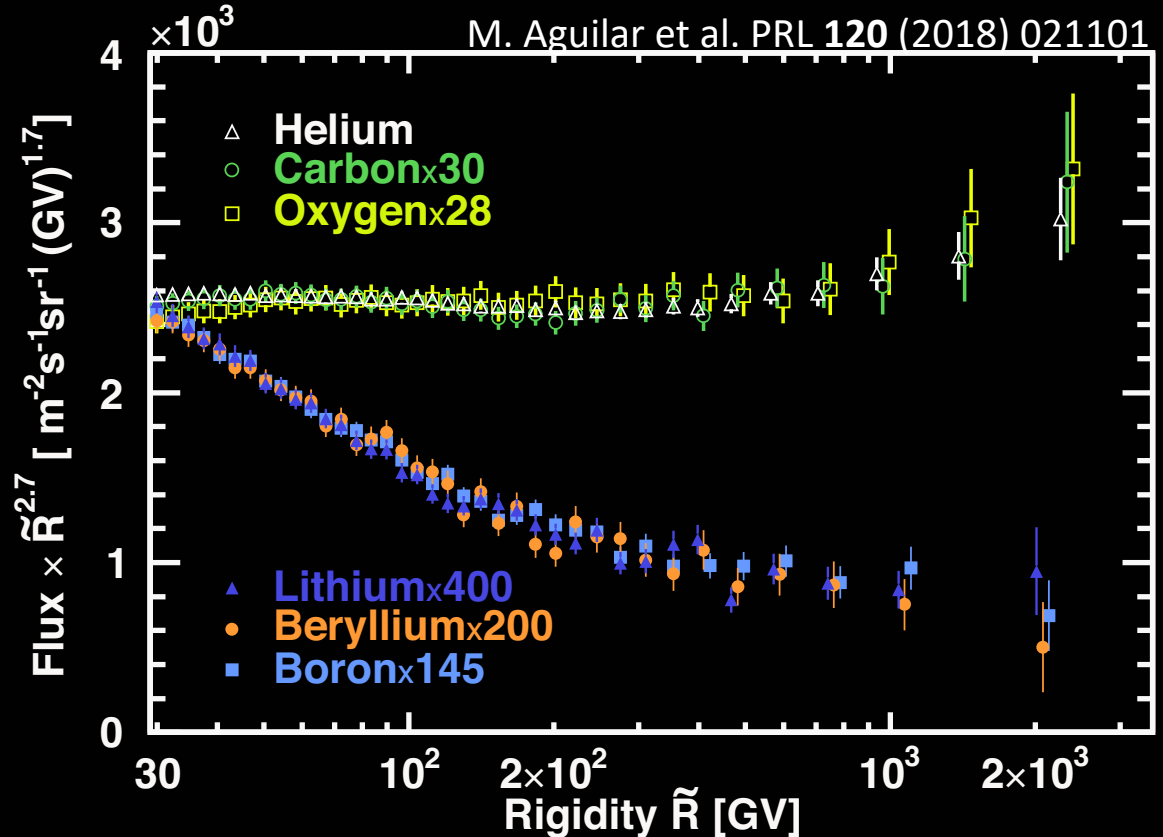


- 20
- AMS01(1998/06)
 - ATIC02(2003/01)
 - ▲ Balloon(1970/09+1971/05)
 - ▼ Balloon(1970/11)
 - Balloon(1976/05)
 - Balloon(1979/06)
 - △ Balloon(1991/09)
 - ◆ BESS-PolarI(2004/12)
 - ⊕ BESS-PolarII(2007/12-2008/01)
 - ★ BESS-TeV(2002/08)
 - ★ BESS98(1998/07)
 - CAPRICE94(1994/08)
 - CAPRICE98(1998/05)
 - ▲ CREAM-I(2004/12-2005/01)
 - ▼ IMAX92(1992/07)
 - LEAP(1987/08)
 - MASS91(1991/09)
 - △ PAMELA(2006/07-2008/12)
 - ◇ PAMELA-CALO(2006/06-2010/01)
 - ◆ RICH-II(1997/10)
 - ★ SOKOL(1984/03-1986/01)
 - ATIC02(2003/01)
 - Balloon(1971/09+1972/10)
 - ▲ Balloon(1972/10)
 - ▼ Balloon(1976/10)
 - Balloon(1991/09)
 - CREAM-II(2005/12-2006/01)
 - △ CRN-Spacelab2(1985/07-1985/08)
 - ◇ HEAO3-C2(1979/10-1980/06)
 - ⊕ PAMELA(2006/07-2008/03)
 - ★ TRACER06(2006/07)
 - ATIC02(2003/01)
 - Balloon(1971/09+1972/10)
 - ▲ Balloon(1972/10)
 - ▼ Balloon(1976/10)
 - Balloon(1991/09)
 - CREAM-II(2005/12-2006/01)
 - △ CRN-Spacelab2(1985/07-1985/08)
 - ◇ HEAO3-C2(1979/10-1980/06)
 - ⊕ TRACER06(2006/07)
 - ★ TRACER99
- 35
- 17

AMS-02 Primaries e Secondaries

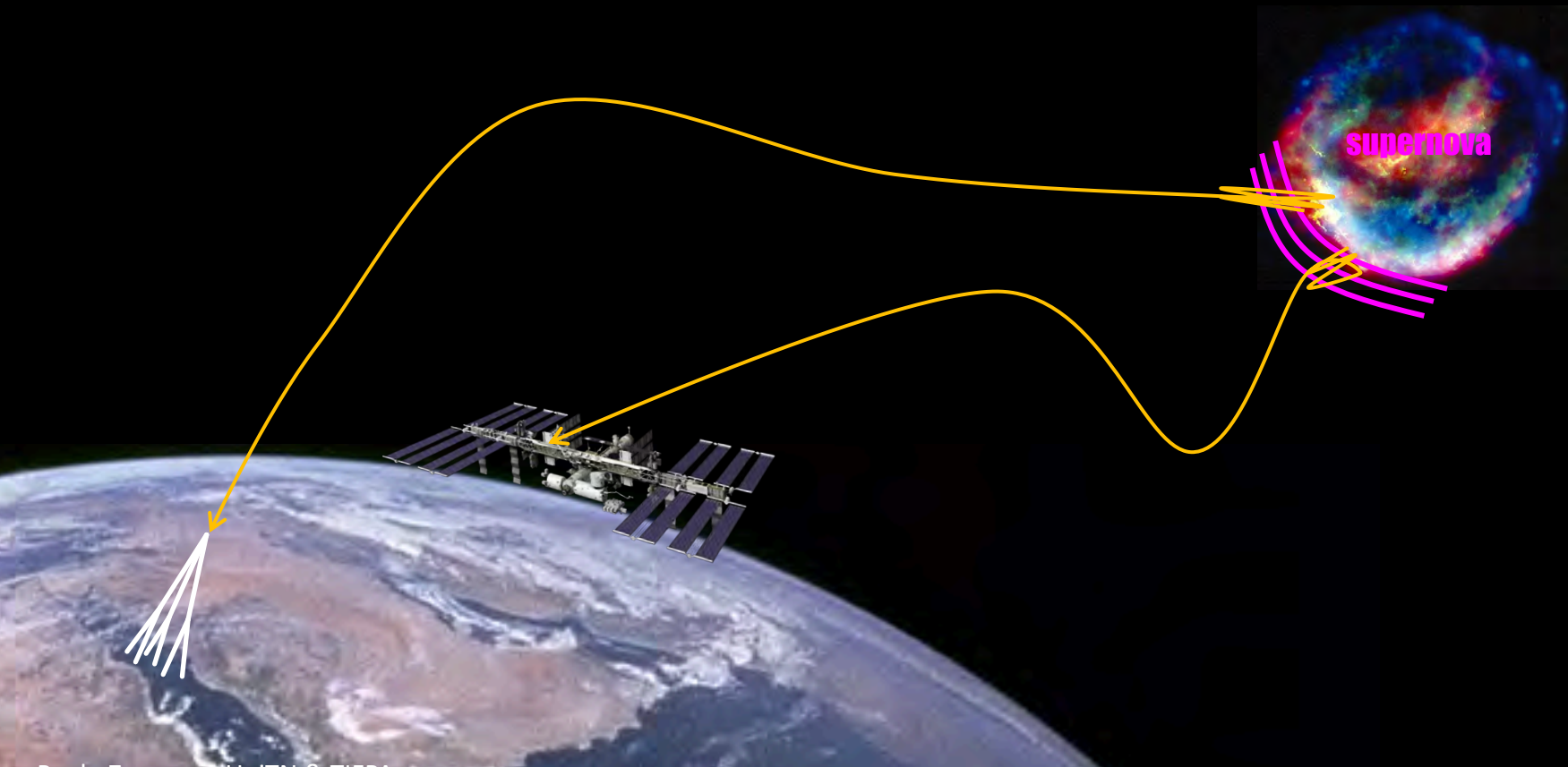
Helium
Carbon (×30)
Oxygen (×28)

Lithium (×200)
Beryllium (×400)
Boron (×145)

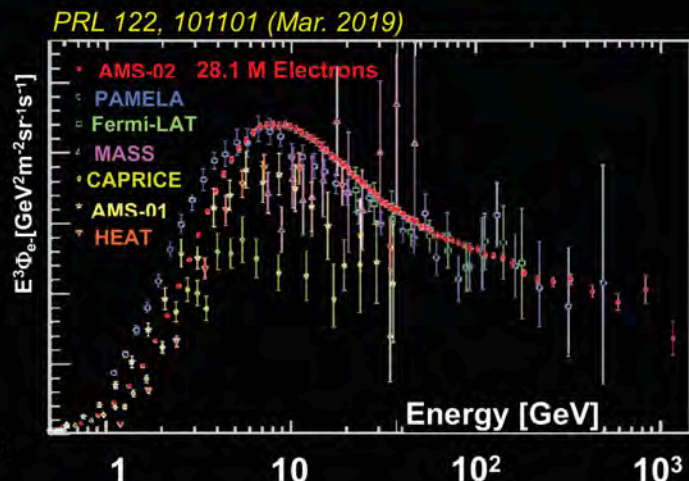
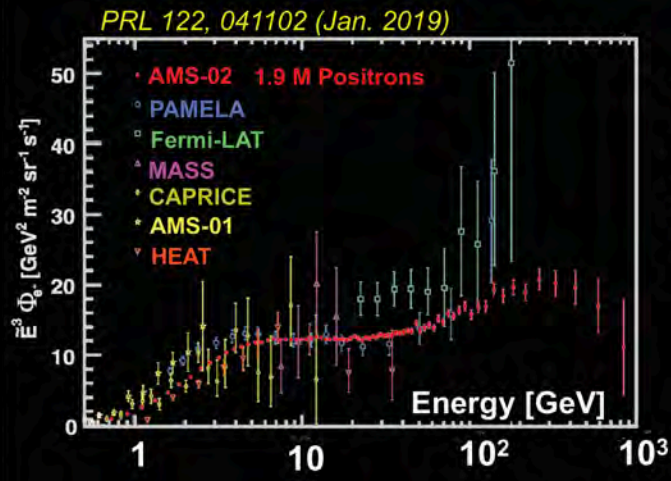


Matter and Anti-Matter in Cosmic Rays

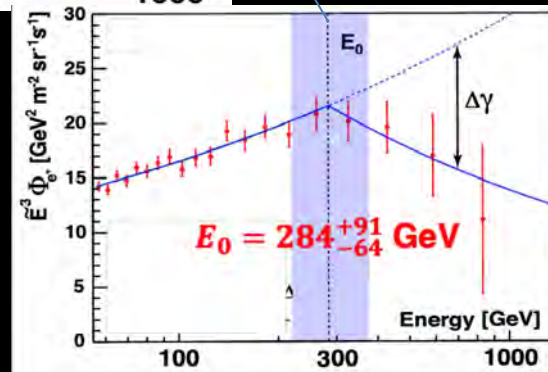
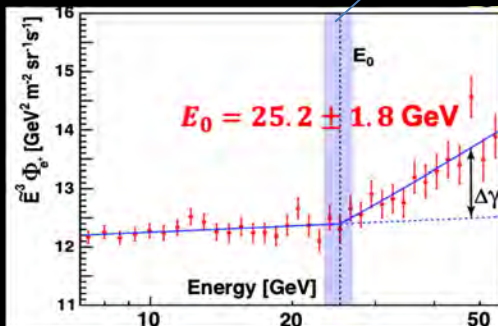
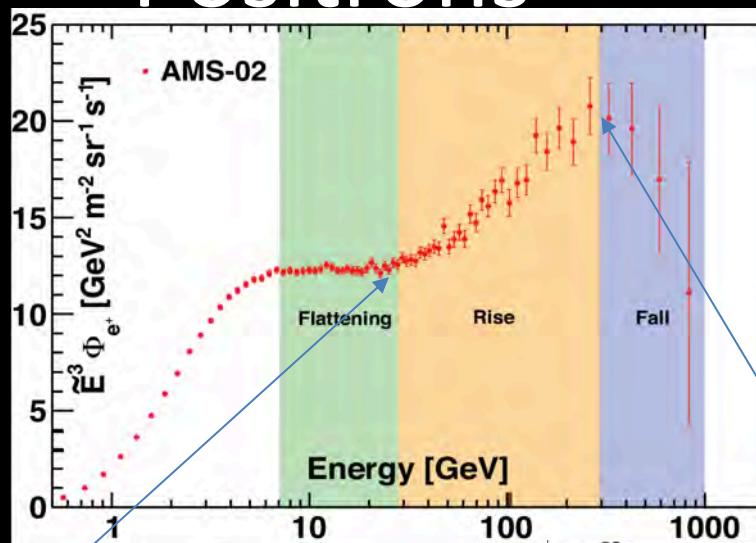
22



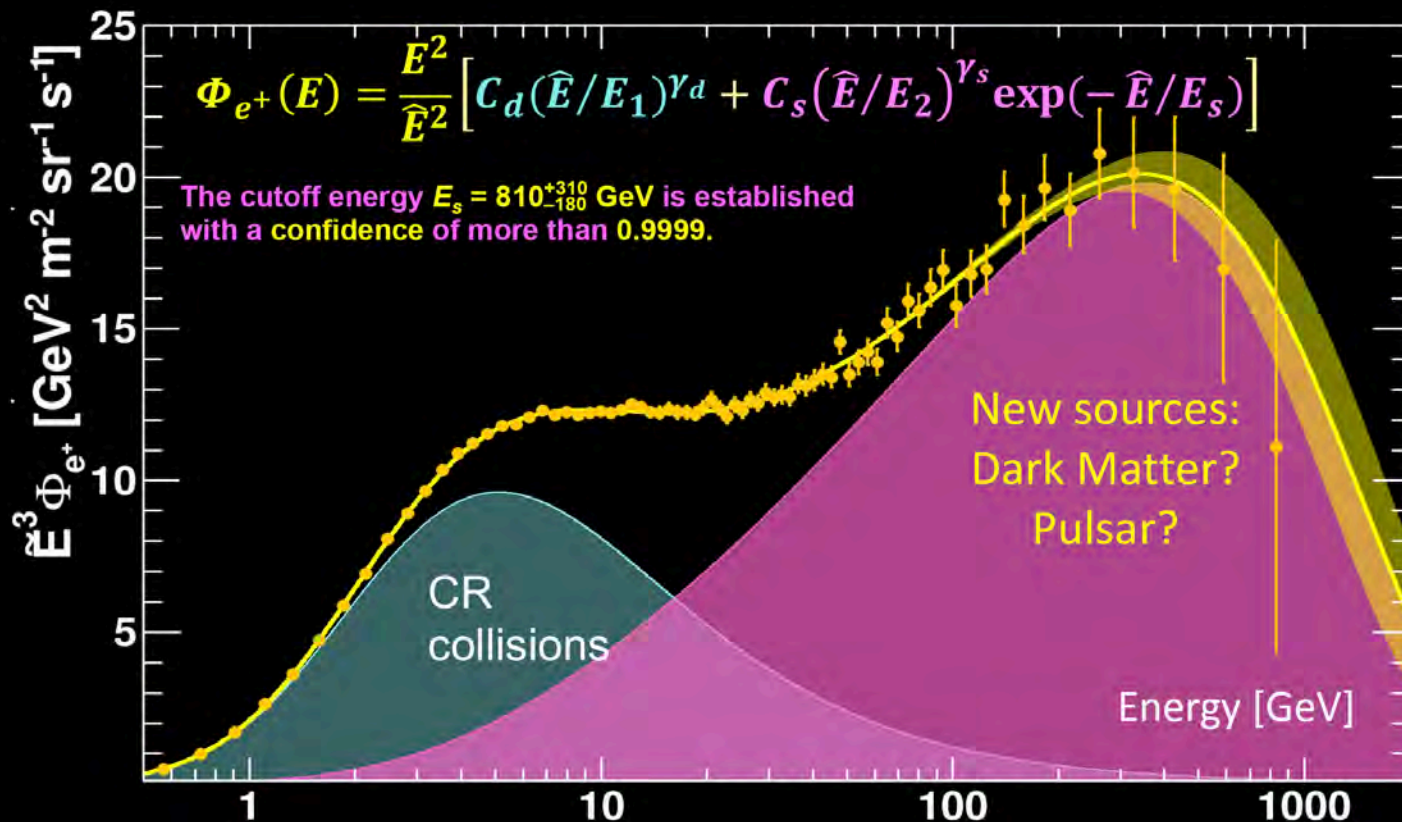
Positron and Electron fluxes



Positrons



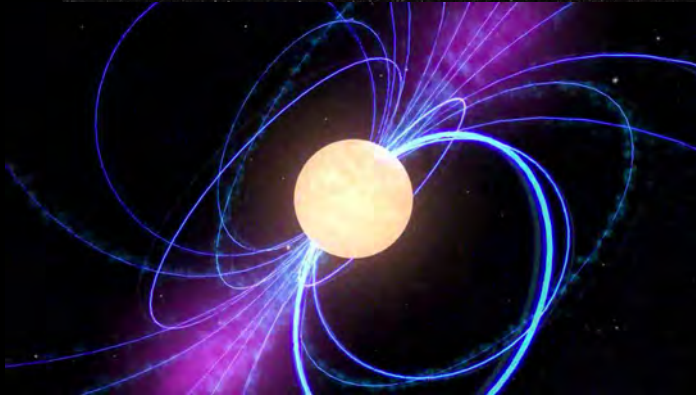
Positron Flux: two components



The origin of positron excess ?

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Pulsars



**Secondary
Production**

supernova

primaries

**Inter Stellar
medium**

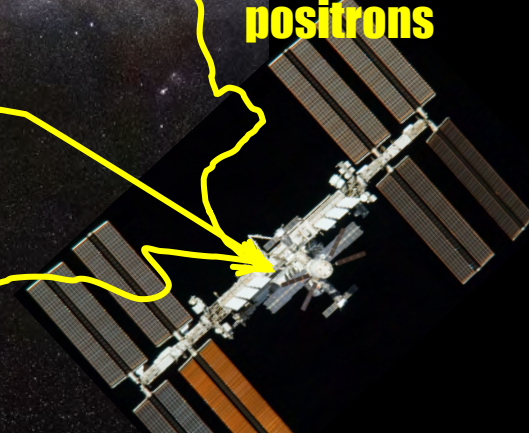
positrons

Dark Matter

electrons

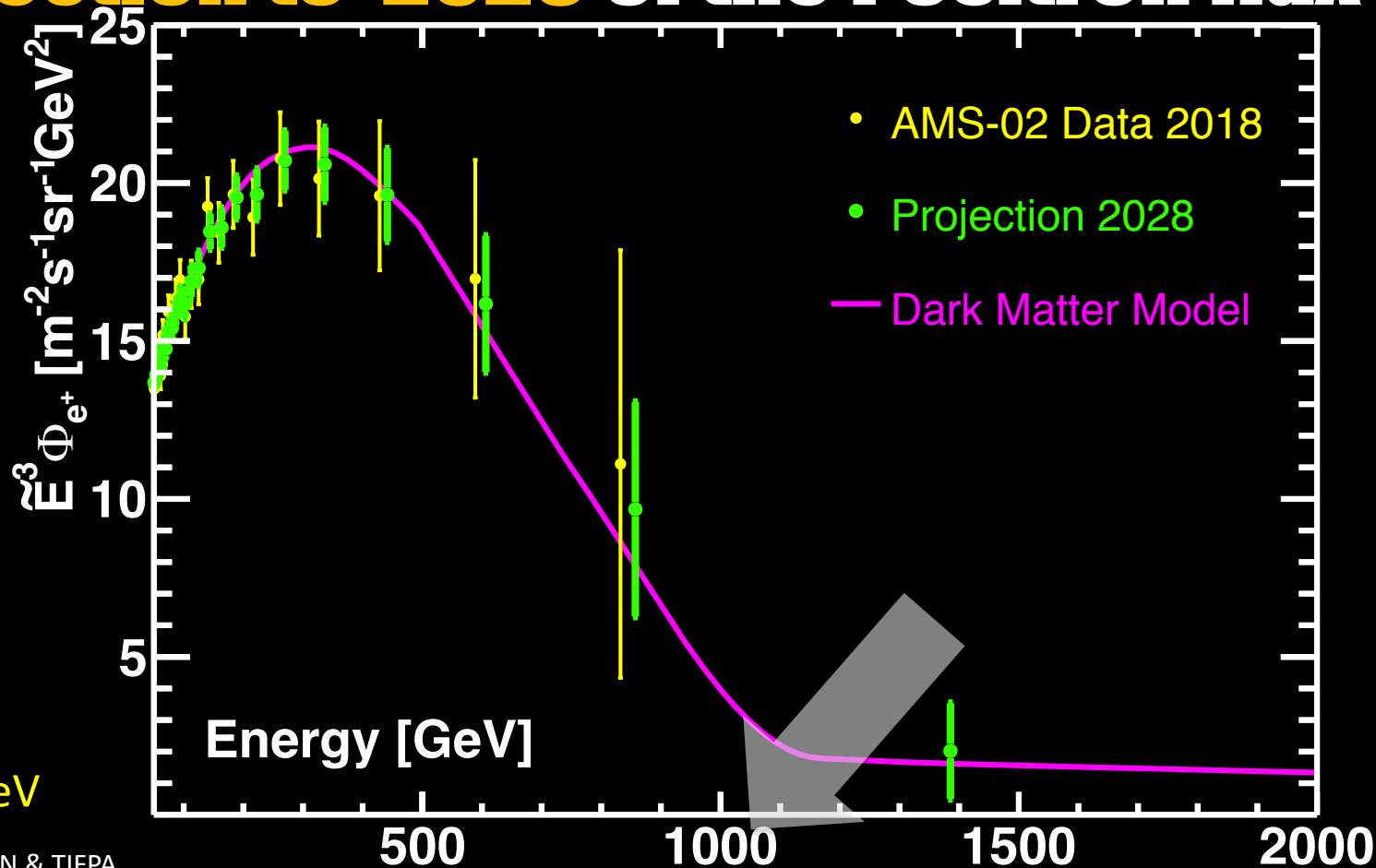
positrons

Dark matter



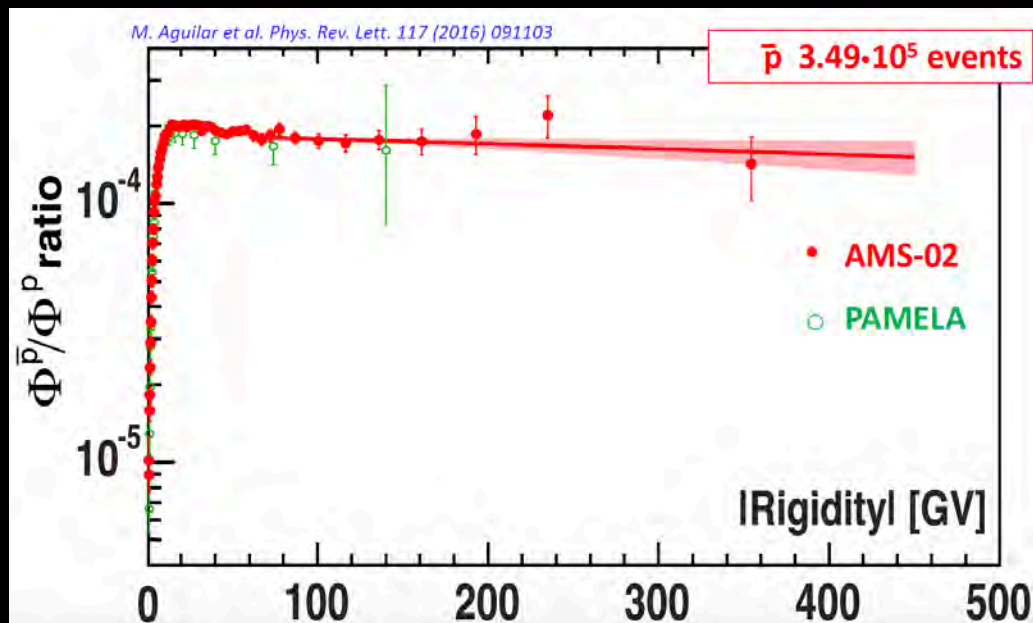
Projection to 2028 of the Positron flux

27

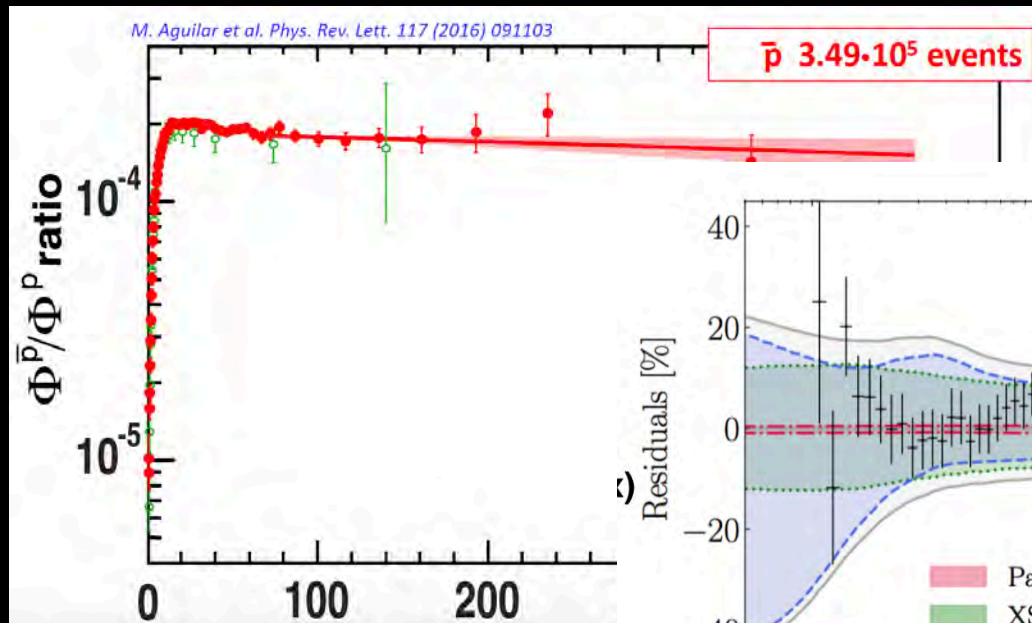


Max ~ 284 GeV

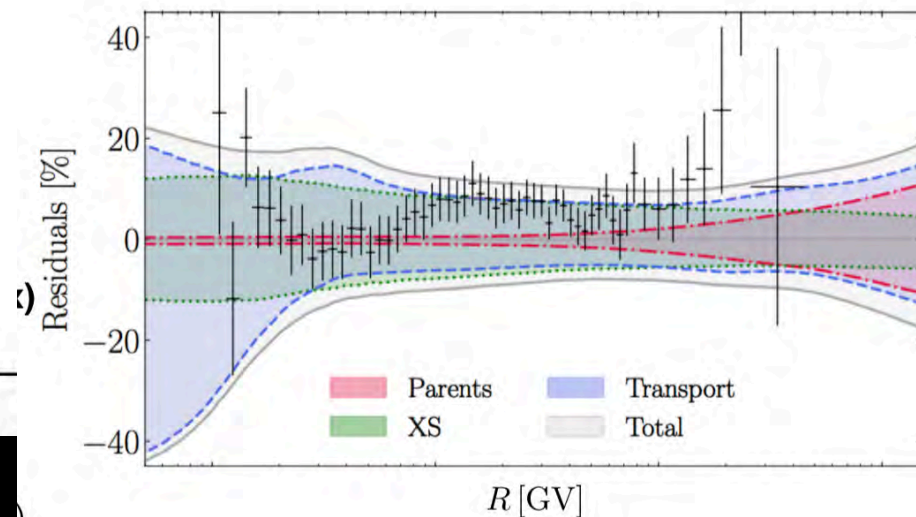
AMS anti-proton to proton ratio



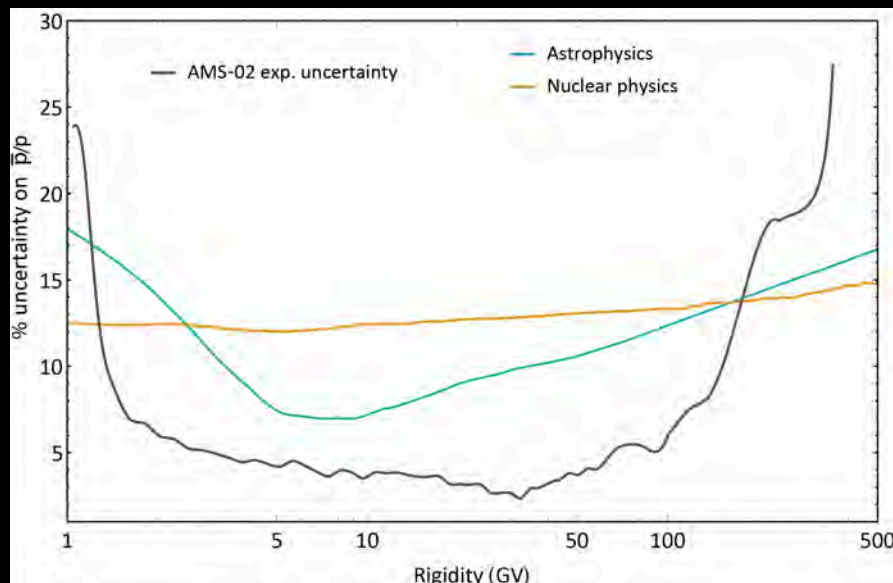
AMS anti-proton to proton ratio



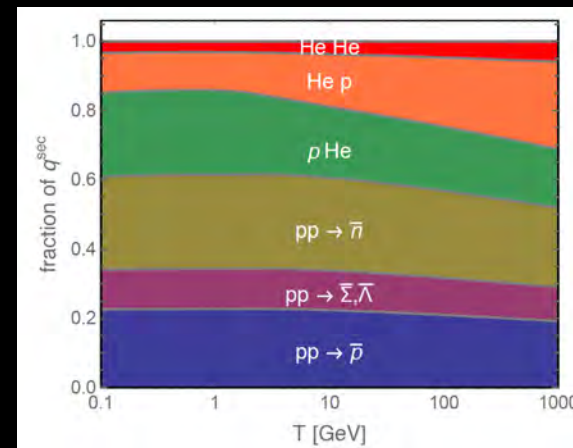
M. Boudaud –ICRC 2019



Prediction Uncertainties



Nicolò Masi, INFN-Bo



AMBER @ CERN

EUROPEAN ORGANIZATION FOR NUCLEAR RESEARCH



CERN-SPSC-2019-003
SPSC-I-250
January 28, 2019

Letter of Intent:

A New QCD facility at the M2 beam line of the CERN SPS*

COMPASS++/AMBER

B. Adams^{1,2}, C.A. Aidala³, R. Akhmetzhanov⁴, G.D. Alexeev¹⁰, M.G. Alexeev¹¹, A. Amoroso^{10,11},
V. Andrieu^{4,5}, N.V. Anisimov¹³, V. Anosov¹², A. Antoshkin¹⁴, K. Augsten^{15,16}, W. Augustyniak¹⁶

RESEARCH



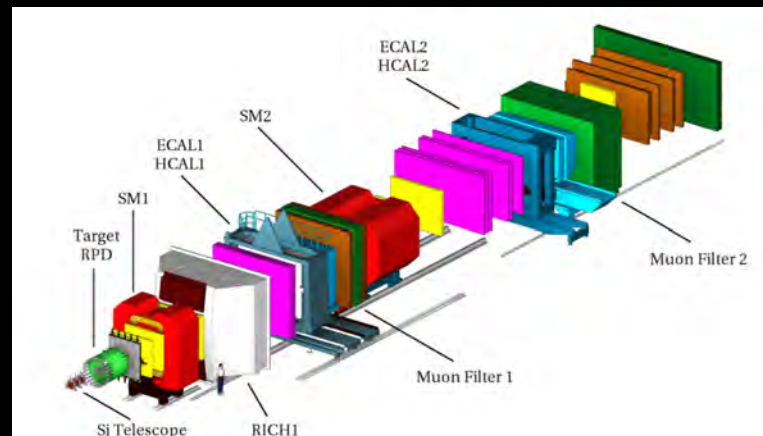
RN-SPSC-2019-022
SPSC-P-360
October 14, 2019

Proposal for Measurements at the M2 beam line of the CERN SPS

- Phase-1 -

COMPASS++/AMBER

COMPASS spectrometer

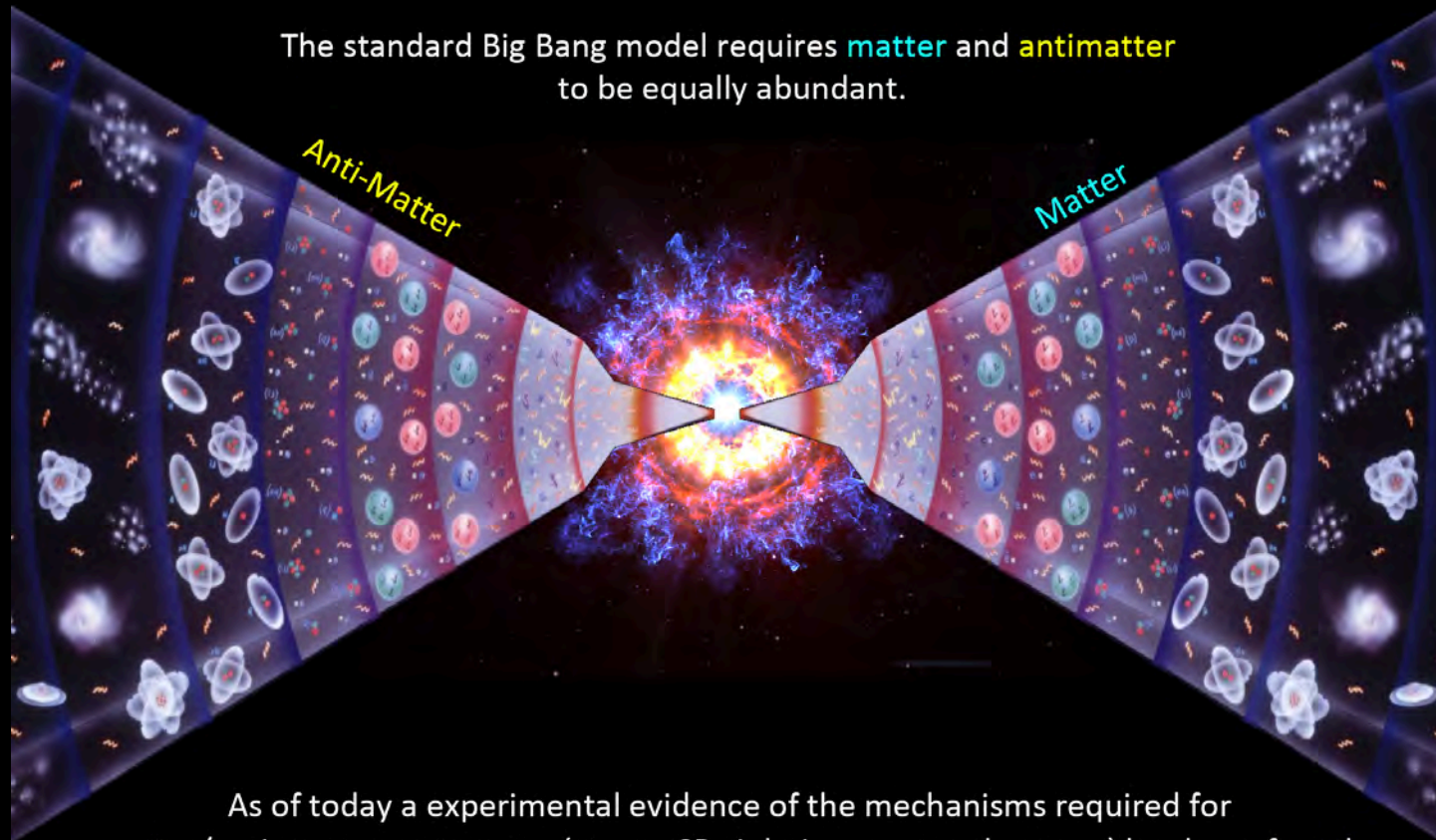


2022-2023 @ CERN SPS

Heavy Anti-Matter Search

33

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The standard Big Bang model requires **matter** and **antimatter** to be equally abundant.

As of today a experimental evidence of the mechanisms required for matter/anti-matter asymmetry (strong CP violation, proton decay, ...) has been found.

Neither has a single anti-nucleus been seen in cosmic rays.

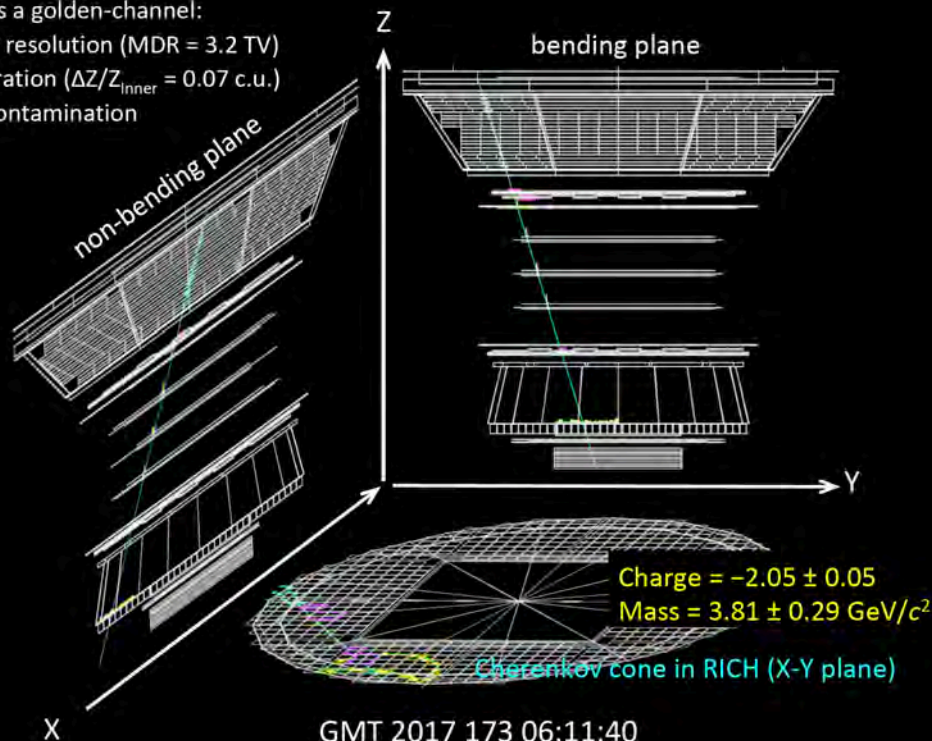
Anti-Helium a candidate event

33

Anti-helium is a *golden-channel*

Anti-helium is a golden-channel:

- Best rigidity resolution ($\text{MDR} = 3.2 \text{ TV}$)
- Best Z separation ($\Delta Z/Z_{\text{inner}} = 0.07 \text{ c.u.}$)
- No p, K, π contamination



Anti-Helium Search Status

Currently, AMS observed 8 anti-helium candidates (mass region from 0-10 GeV/c^2) with rigidity < 50 GV with respect to a sample of 700 million helium events selected.

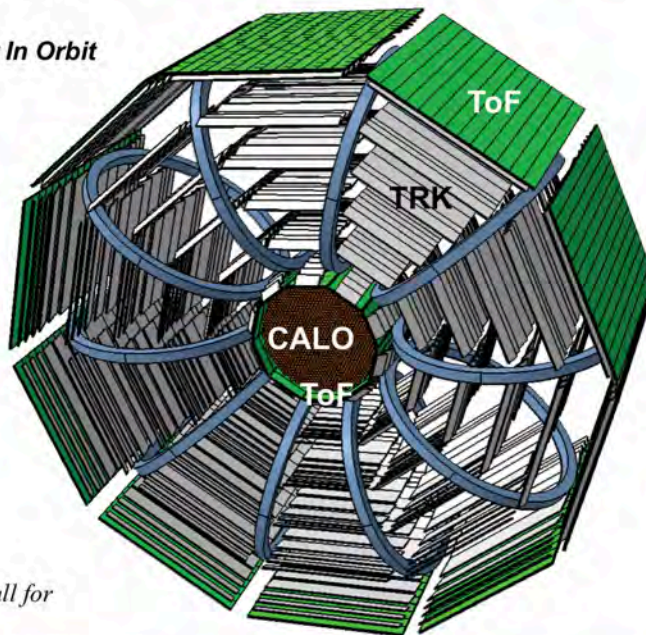
The rate in AMS of antihelium candidates is less than 1 in 100 million helium.

At this extremely low rate, more data (**through the lifetime of the ISS**) is required to further check the origin of these events.



What's next ?

**High Precision Particle Astrophysics as a
New Window on the Universe**
*with an Antimatter Large Acceptance Detector In Orbit
(ALADInO)*



*A White Paper submitted in response to ESA's Call for
the VOYAGE 2050 long-term plan*

<https://www.cosmos.esa.int/web/voyage-2050/white-papers>

https://www.cosmos.esa.int/documents/1866264/3219248/BattistonR_ALADINO_PROPOSAL_20190805_v1.pdf

Summary

- The recent observations point consistently to an Universe filled of dark-matter
- A lot of questions are open and the natural laboratory of the space is a place to search for answers.
- We live in a exciting age where the large knowledge we built, disclose to us an uncharted new landscape.
- Cosmic rays are one of the open window to the Universe and they have the potential to reveal new physics
- Current and future experiments measuring cosmic rays may help in writing the map of this new territory.



<https://AMS02.space>



AMS-02 represents a big success and shows, the scientific and technologic excellence achieved by INFN, ASI and the Italian Universities.



Please visit